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COMPARATIVE STABLE ISOTOPES STUDY OF THE MAGNESITE DEPOSITS OF GREECE

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ABSTRACT

The magnesite deposits of North Evia, Lesvos and Chalkidiki, which are hosted by Alpine type ultramafic rocks, have been studied from an isotopic point of view and are compared with similar magnesite deposits occurring in other countries of central-southeastern Europe and Turkey.

Magnesites from N. Evia, Lesvos and Chalkidiki have similar isotopic compositions. They have $\delta^{13}C_{\text{PDB}}$ values ranging from -8.6 % to -14.8 % and $\delta^{18}O_{\text{SMOW}}$ values ranging from 24.0 % to 30.1 %, which indicate a rather similar formation process for all three areas.

The studied Greek magnesite deposits have similar isotopic compositions with vein and stockwork type magnesite deposits from: Turkey (Koyakci Tepe, vein and stockwork types; Helvacibaba, stockwork type), Poland (vein type), and Former Yugoslavia (vein type).

There is strong isotopic evidence that meteoric water is involved in the formation of the stockwork and vein type magnesite deposits of all four countries, including Greece. There is also isotopic evidence for the organic origin of carbon participating in the carbonates (magnesites) of these deposits.

1 INTRODUCTION

North Evia and Chalkidiki are the major magnesite producing areas of Greece. There have been numerous studies on the magnesite deposits of Evia, Lesvos, Chalkidiki and similar magnesite deposits all over the world hosted by ultramafic rocks (Petraschech 1964, Capedri & Rossi 1973, Dabitzias 1980). Two independent studies on the magnesite deposits of Chalkidiki (Vavdos), the first by Dabitzias (1981) and the second by Burgath et al. (1981), suggested different origin for the same magnesite deposits, as hydrothermal or as weathering products, respectively. These studies reveal a general controversy on the origin of magnesite problem becomes more complicated as Illic (1968) suggested that medium- to large-sized veins of magnesite, which occur at deeper level, are hydrothermal while isolated near-surface nodules and subspherical bodies of magnesite are products of weathering. Suggested temperatures for the precipitation of magnesite range from 10 to 400°C. Pohl (1990) discussed and summarized models and trends concerning the genesis of magnesite.

Gartzos (1986, 1990) on the basis of geochemical - isotopic data and vein petrography studies suggested that the magnesite deposits of N. Evia were formed from mixing of descending meteoric water with ascending CO_2 at temperatures <60°C. Barnes et al. (1986) presented interesting stable isotopes data concerning the origin of thermal and cold meteoric waters and dissolved bicarbonates occurring in the Peri - Aegean region (Fig. 1 & 2, Tab. 1). These data have been used in the present study in order to investigate the nature of the aqueous solutions, which were involved in the genesis of the studied magnesite deposits.

The last decade a lot of magnesite stable isotopes data have been accumulated, which could support a comparative study for a better understanding of the origin of carbon and involved water.

In the present study, the magnesite deposits of N. Evia, Lesvos, and Chalkidiki (Fig.1) are studied and compared from a stable isotopes point of view. They are also compared with similar

magnesite deposits from other countries. The origin of carbon and the type of the waters (aqueous fluids) involved are inferred from stable isotopes data.



Figure 1. Map of Greece showing the magnesite deposits of Evia, Lesvos and Chalkidiki plus sampling locations of the reported cold meteoric waters and thermal waters.

2 GEOLOGICAL SETTING

The areas of N. Evia, Lesvos and Chalkidiki are highly faulted due to extentional tectonics prevailing in the Aegean and peri-Aegean region during the last 13 million years (Le Pichon & Angelier 1979). The magnesite deposits occur in ultramafic rocks and are closely associated with these faults. The most important deposits are hosted by non or partly serpentinized peridotites, which usually have been altered to brown serpentinite (a rock composed mainly of carbonates, quartz, limonite and serpentine) at the sites of intense mineralization or towards the contacts of some veins. The known deposits are restricted to shallow depths (maximum depth is about 250 m). In all three areas there is a considerable amount of ophiolites, which have been overthrusted on Mesozoic platform carbonates. They are dismembered ophiolites with missing upper units. Sheeted dyke complexes, pillow lavas, and sediments have been either sheared off or weathered out (Guernet 1971, Caperdi 1974, Dabitzias 1980, Gartzos 1986, Gartzos et al. 1994, Serelis & Gartzos 1994).

3 TYPE OF MINERALIZATION

Magnesite is cryptocrystalline with crystals 1-5 μ m large. It is usually snow white, shows conchoidal fracture and resembles unglazed porcelain. On the basis of the type of mineralization three main types of magnesite deposits were distinguished: a. Nodular aggregates of magnesite replacing the host rock. Very often, coalesced nodules form irregular to roughly subspherical bodies of magnesite varying from a few cm to a few meters in diameter. They usually occur at higher horizons (nearsurface).

b. Fracture - filling magnesite, which occurs as veins along faults, joints and other fractures. The veins, which occur at the upper parts of the deposits (near-surface) are thin and usually form dense networks (stockwork - type deposits). Medium sized veins (5-25 cm) and larger ones (25-100 cm) occur at deeper levels. The majority of the veins have nodular structure.

c. Lens - like bodies of magnesite (1 to 4 m wide and 5 to 12 m long) formed by coalescence of nodules in dilatation zones of pinch - and - swell structures. They are located at deeper horizons.

4 MATERIALS AND METHODS

Magnesites and dolomites associated with the peridotites and serpentinites of North Evia, Lesvos and Chalkidiki were sampled and analyzed for their $\bar{\delta}^{13}C$ and $\bar{\delta}^{18}O$ isotopic compositions. The samples of dolomite are from Evia where dolomite has been formed by replacement of magnesite and serpentine.

All samples were reacted with phosphoric acid at 50^oC. The isotope ratios were determined on a VG Micromass 903 triple collector mass spectrometer at the stable isotopes laboratory of the ETH in Zurich. The samples were measured against the Carrara Marble laboratory standard. Carbon isotopes ratios of the carbonates are reported in permil deviations relative to the Chicago PDB standard, whereas oxygen isotopes ratios are given relative to SMOW.

The oxygen and carbon isotope compositions of the analyzed carbonates are presented in table 1 and are shown in figure 2. Identical results were obtained from samples SOT.1* and SOT.1** which represent two independent analyses of powder prepared from hand specimen SOT.1 and reacted separately.

5 ISOTOPIC RESULTS

The δ -values (vs SMOW) of the oxygen isotopes of the magnesites from N. Evia, Lesvos and Chalkidiki range from 24.0 to 30.1 ‰, and carbon isotopes δ -values (vs PDB) vary from -8.6 to -14.8 ‰ (Tab. 1). These δ -values correspond to magnesite values, which O'Neil & Barnes (1971) consider to represent isotopic equilibrium with their formation waters as they plot in the equilibrium field defined by them (Fig. 2B). The field has been defined on the basis of isotopic data from carbonates associated with ultramafic rocks occurring along the central and northern Coast Ranges of California (Barnes & O'Neil 1971, Barnes et al. 1973).

Nr	Sample	Phase	δ ¹⁸ Ο	δ ¹³ C	Nr	Sample	Phase	δ ¹⁸ Ο	δ ¹³ C
			SMOW	PDB				SMOW	PDB
N.Evia				15	MG*	magn.	28.6	-10.9	
1	Sot.1*	magn.	27.7	-10.6	16	MG**	magn.	27.1	-9.3
2	Sot.1**	magn.	27.7	-10.6	17	Limni	magn.	27.4	-8.6
3	Sot. 1B	magn.	29.1	-10.4	18	Kak.53/11x	dolom.	25.6	-12.2
4	Kak.4	magn.	26.4	-10.6	19	Tr.6/4x	magn.	26.7	-11.9
5	Kak.53	magn.	28.2	-11.1	20	Tr.7/3x	dolom.	26.4	-10.6
6	Kak.2A	magn.	25.4	-10.8	Lesvos				
7	Pars.20	magn.	28.5	-10.0	21	Lv.1	magn.	26.0	-13.3
8	Pars.23	magn.	29.8	-9.9	22	Lv.2	magn.	27.7	-12.9
9	Pars.22C	magn.	30.1	-9.9	23	Lv.3	magn.	27.2	-12.5
10	Pars.21	magn.	29.8	-9.7	24	Lv.4	magn.	24.0	-13.9
11	Ger.11	magn.	27.7	-11.5	Chalkidiki				
12	Pap.6	magn.	28.8	-12.3	25	Ch.1	magn.	28.6	-14.5
13	Pink	magn.	28.7	-13.3	26	Ch.2	magn.	27.4	-14.8
14	MG	magn.	27.2	-9.9	27	Ch.4	magn.	28.0	-14.3

Table 1. $\delta^{18}O_{SMOW}$ and $\delta^{13}C_{PDB}$ values of magnesites and dolomites from N. Evia, Lesvos and Chalkidiki.



Figure 2. A: Plot of the peri-Aegean thermal waters (•) and cold meteoric waters (•) in the δD vs $\delta^{18}O$ diagram. B: Correlation plot between carbon and oxygen isotopes contained in the HCO₃⁻ of thermal (o) and cold (□) waters and in the magnesites of N. Evia (▲), Lesvos (★) and Chalkidiki (*). Symbols as in figure 1.

6 ORIGIN OF WATER AND CARBON

6.1 Thermal and cold waters of the Aegean and Peri-Aegean area

It was mentioned earlier that magnesite is a common fracture filling phase in the ultramafic rocks of N. Evia, Lesvos and Chalkidiki. These host rocks show a typical "brown alteration". These observations make clear that magnesite has been formed by waters circulating through the ultramafic rocks.

Waters occurring in mineralized and nonmineralized ultramafic rocks of Evia, Lesvos and Chalkidiki, and other localities of the peri-Aegean area have been studied by Papastamataki & Leonis (1982), Barnes et al. (1986), Gartzos (1986), Gartzos & Stamatis (1996). These waters have been plotted on the δD vs δ^{18} O diagram shown in figure 2A. The cold meteoric waters are samples from: (a) meteorological stations and (b) cold springs and rivers which are supposed to represent

homogenized local meteoric waters. The thermal waters include samples from: (a) thermal springs and (b) exploration boreholes in hydrothermal fields.

The cold meteoric waters plot along the meteoric water line $\delta D=8\delta O^{18}+10$ (Craig 1963). Their $\delta^{18}O_{SMOW}$ permit values range from -9.5 to -4.3 (mean value = -7.2). The magnesites from Evia, Lesvos and Chalkidiki, which are shown in figure 2B have $\delta^{18}O_{SMOW}$ permit values ranging from 24.0 to 30.1 (mean value = 27.7).

Assuming equilibrium fractionation between magnesite and meteoric water, a fractionation of about 35 is estimated from figure 2. Similar fractionation of about 34 at 12° C (mean temperature of cold meteoric waters) has been obtained from the equation of O'Neil & Barnes (1971). Unfortunately experimental isotopic data for the system MgCO₃ - H₂O are lacking. However, the compiled 1000. In. α vs T curves (Friedman & O'Neil, 1977) for the fractionation of O¹⁸ between alkaline earth carbonates and water can serve as a crude guide for the magnesite - water system. This is justified by the fact that the O¹⁸ fractionation curves (1000.In. α vs T) of all the experimentally determined carbonate - water systems (BaCO₃ -, SrCO₃ -, CaCO₃ - and CaMg(CO₃)₂ - H₂O) show parallel trend and plot close to each other. Thus, in analogy, the fractionation curve of magnesite is expected to plot very close to the curve of dolomite in a parallel trend. Assuming that the above approximation is correct a fractionation in the range from 33 to 36 is predicted for the magnesite - water fractionation estimated from both, figure 2 and the equation of O'Neil & Barnes (1971).

The thermal waters (mean temperature = 50° C) have been also plotted in figure 2. On the basis of hydrochemical analyses and isotopic data the following were suggested concerning the origin of the thermal waters (Barnes et al. 1986; Gartzos & Stamatis 1996). Sample 22 is a low TDS (<1000 ppm) thermal water, which contains 20 ppm Cl⁻. It is heated local meteoric water. Samples 19, 20, and 24-27 are high TDS (>1000 ppm) thermal waters with Cl⁻ ranging from 970 ppm to 4700 ppm. They were generated by mixing of ascending hot sea water with descending local meteoric water. The contribution of seawater ranges from 10% to 25%. Samples 18, 21a, 21b and 23 are also high TDS thermal waters but with Cl⁻ ranging from 18970 ppm to 21985 ppm. The contribution o seawater ranges from 92% to 100%. These waters plot very close to SMOW (Fig. 2A).

Thermal water $\delta^{18}O_{SMOW}$ permil values range from -9.5 to 1.0 (mean value = -4.0). Figure 2A shows that the thermal waters with low $\delta^{18}O_{SMOW}$ values plot in the cold meteoric water field. However, considering the mentioned curves showing the fractionation of O¹⁸ between alkaline earth carbonates and water at various temperatures (Friedman & O'Neil, 1977), as well as the above mentioned equation of O'Neil & Barnes (1971), a drastic fractionation decrease with increasing temperature is expected. At 50°C (mean temperature of thermal waters at the site of discharge - earth surface) a fractionation of about 26 is expected. However, depth temperatures >100°C have been reported by Papastamataki & Leonis (1982) for the high δ^{18} O thermal waters. At 100°C fractionation is expected even smaller, about 22. Magnesites formed from these thermal waters are expected to have different $\delta^{18}O_{SMOW}$ values from those observed in the magnesites of N. Evia, Lesvos and Chalkidiki. They are expected to range from 16.5 to 23.0 permil, assuming stronger fractionation between magnesite and low $\delta^{18}O$ water (-9.5 + 26.0 = 16,5 $\delta^{18}O_{SMOW}$ permil) and weaker fractionation between magnesite and high $\delta^{18}O$ water of higher temperature (1.0 + 22.0 = 23.0 $\delta^{18}O_{SMOW}$ permil).

All the above relations and indications suggest that a meteoric origin for the water involved in the genesis of magnesite should be considered seriously.

6.2 HCO₃⁻ in thermal and cold waters of the peri-Aegean region.

Cold waters occurring in the peri-Aegean area contain bicarbonates with $\delta^{13}C_{\text{PDB}}$ values from -8.0 to -15.5 ‰ and $\delta^{18}O_{\text{SMOW}}$ from 24.5 to about 30.0 ‰, which are in the same range with $\delta^{13}C_{\text{PDB}}$ and $\delta^{18}O_{\text{SMOW}}$ values observed in magnesites from N. Evia, Lesvos and Chalkidiki (Fig. 2B). Thermal water bicarbonates have similar $\delta^{18}O_{\text{SMOW}}$ values, but they can be distinguished by the different $\delta^{13}C_{\text{PDB}}$ values.

Barnes et al. (1986) suggest that δ^{18} O of HCO₃⁻ is controlled by δ^{18} O of H₂O in both thermal and cold waters. This is expected because the oxygen reservoir of the water is very large and controls the small oxygen reservoir of the dissolved HCO₃⁻. On the contrary, H₂O itself has no carbon and the aqueous species of carbon form a small carbon reservoir in the aqueous solutions. Varia-

200

tions in the carbon sources themselves or in their relative contribution will affect the C¹³/C¹² ratios of both, the aqueous species of carbon and the formed magnesites. Such seems to be the case in our study area. Consequently, the carbon isotopes only can serve as indicators of sources. Fractionation of C¹³ between alkaline earth carbonates and HCO₃⁻ is expected to be very low

Fractionation of C¹³ between alkaline earth carbonates and HCO₃⁻ is expected to be very low at cold-water temperatures. Reported relevant experimental data by Emrich et al. (1970) indicate the following fractionation factors for the system calcite - HCO₃⁻: 1.001 at 12 °C, 1.003 at 50°C and 1.0042 at 100°C. Consequently, cold meteoric waters are expected to precipitate magnesite with $\delta^{13}C_{PDB}$ values similar or very close to those observed in the dissolved bicarbonates.

On the contrary, thermal water bicarbonates have much higher $\delta^{13}C_{PDB}$ values (Fig. 2B). In addition, the magnesite - HCO₃ fractionation is expected to be higher due to the higher temperatures of the thermal waters. Consequently, magnesites formed from these waters are expected to have relatively high $\delta^{13}C_{PDB}$ values, about 5 %.

Concerning the negative $\delta^{13}C_{PDB}$ values of the bicarbonates Barnes et al. (1986) suggest that it is due to contribution of organic CO₂ from the soil, where micro organisms and the roots of plants produce CO₂.

7 COMPARISON WITH MAGNESITE DEPOSITS FROM OTHER COUNTRIES

Our stable isotope data plus data from former Yugoslavia (Fallick et al. 1991), Turkey (Zedef et al. 2000) and Poland (Jedrysek & Halas 1990) are shown and correlated in figure 3.



Figure 3. $\delta^{18}O_{SMOW}$ versus $\delta^{13}C_{PDB}$ diagram with data from Greece, Former Yugoslavia, Poland and Turkey, Greek vein magnesites: field is defined by solid line (square).

Polish vein magnesites: field is defined by dotted line.

Former Yugoslavia vein magnesites: field is defined by dashed line with three dots between.

Former Yugoslavia Lacustrine magnesite (Bella Stena): field is defined by dashed line.

Former Yugoslavia intermediate type: field is defined by dashed line with one dot inbetween.

Turkish magnesites are plotted in the diagram (see KEY for localities-symbols).

The Greek vein magnesites overlap and correlate very well with vein-type magnesites from all three other countries. The field of vein type magnesites is located between the lacustrine magnesite field, shown in the upper right part of the diagram, and the Seydisehir carbonates occurring in meta argillites (rich in organic carbon). According to Zedef et al. (2000) the Seydisehir carbonates and

the lacustrine magnesites form the two end members of a fluid mixing line, which connects them. In the middle of this mixing line is the field of the vein type magnesite. Zedef et al. (2000) suggest that the negative carbon values are due to the contribution of organic CO_2 from the decarboxylation of organic material occurring in the sediments. According to Irwin et al. (1977) this decarboxylation reaction starts at about 70°C.

8 CONCLUSIONS

The oxygen isotope compositions of the magnesites from N.Evia, Lesvos and Chalkidiki range from 24.0 to 30.1 ‰, and carbon isotope compositions vary from -8.63 to -14.85 ‰. Similar isotopic compositions have been reported for similar vein and stockwork type magnesite deposits from: Turkey (Koyakci Tepe, vein and stockwork types; Helvacibaba, stockwork type), Poland (vein type), and Former Yugoslavia (vein type).

These compositions correspond to magnesite values, which are considered to represent isotopic equilibrium with their formation waters.

The fractionation between magnesite and cold meteoric water, estimated from the observed $\delta^{18}O_{SMOW}$ values, is about 35. A similar fractionation of about 34 at 12°C (mean temperature of cold meteoric waters occurring in the peri-Aegean area) was obtained using the equation of O'Neil & Barnes (1971). A fractionation of the same order of magnitude (at 12°C) is suggested from consideration of the compiled 1000.ln. α vs T curves (Friedman & O'Neil 1977) for the fractionation of O¹⁸ between alkaline earth carbonates and water. This is strong evidence that meteoric water is involved in the formation of the magnesite.

Cold waters occurring in the peri-Aegean area contain bicarbonates with $\delta^{13}C_{\text{PDB}}$ values from -8.0 to -15.5 % and $\delta^{18}O_{\text{SMOW}}$ from 24.5 to about 30.0 %, which are in the same range with $\delta^{13}C_{\text{PDB}}$ and $\delta^{18}O_{\text{SMOW}}$ values observed in magnesites from N. Evia, Lesvos and Chalkidiki. Thermal water bicarbonates have similar $\delta^{18}O_{\text{SMOW}}$ values with magnesite, but different $\delta^{13}C_{\text{PDB}}$ values. The close isotopic relationship between magnesites and cold-water bicarbonates could indicate that these bicarbonates could be the source of carbon, in such a case carbon is organic from the activities of bacteria and other organisms in the soil. The suggestion of Zedef et al. (2000) for the origin of carbon from decarboxylation reaction at depth is a strong one as well. However, in both cases the interpretation is definitely organic carbon, which is the only carbon source that can give the strongly negative $\delta^{13}C_{\text{PDB}}$ values.

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